Analysis of Background γ-Ray Measured with Shielded HPGe Detector

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1. Introduction

The background γ -ray was measured with shielded HPGe detector. Lead bricks with 10 cm thickness and copper plates with 3 mm thickness were used for the shield. The shield dimension was $1000 \times 1400 \times 1150$ mm³. The background was measured with and without the shield as well as with the copper plates inside of the shield. Also, the X-rays and γ -rays spectra for standard sources, ²⁴¹Am, ¹³³Ba, ¹³⁷Cs, ⁶⁰Co, ¹⁵²Eu, were measured. The γ -ray spectroscopy systems was composed of closed-ended coaxial type HPGe detector, Linear AMP, ADC, and AIM. The relative detection efficiency of the HPGe detector is 18% and the energy resolution is 1.8 keV for the 1.3 MeV γ -ray peak from ⁶⁰Co.

2. Comparison of background spectra under various shield conditions

The measured background spectra are shown in figure 1. Analysis of the background peak areas was performed by HYPERMET code[1].



Figure 1. Background spectra under various shield conditions.

Total count rate of the background was drastically reduced by shield from 187.2 cps to 1.7 cps. Some of the background γ -rays and X-rays peaks is listed in table 1. As for ⁴⁰K, one of the most representative background γ -ray peaks[2], it has decreased by a factor of 200 from 1753.6(3)×10⁻³ cps to 6.3(1)×10⁻³ cps. Count rate of γ -rays from any other background sources such as ²¹⁴Bi, ²²⁸Ac and ²⁰⁸Tl[2,3] has also decreased by a factor of 100.

In the case of 238.6 keV γ -ray emitted from ²¹²Pb, the count rate has decreased by a factor of 20 from 779(7)×10⁻³ cps to 33.4(3)×10⁻³ cps. Some peaks that were from ²¹²Pb and ²³⁵U showed a similar tendency. It is determined that this tendency was shown because of the decreased Compton continuum and lead shield.

Table 1. Major background γ-rays and X-rays.

Nucli- de	Energy (keV)	Count rate(10 ⁻³ cps)		
		Not shielded	Pb- shielded	Pb, Cu- shielded
Pb	74.8*)	394.3(76)	27.8(2)	6.9(2)
Bi	77.1 ^{*)}	258.6(64)	15.1(2)	8.9(3)
²²⁶ Ra	186.0	90.4(31)	10.8(2)	7.7(2)
²¹² Pb	238.6	778.8(67)	33.4(3)	22.4(3)
²¹⁴ Pb	351.9	345.6(20)	6.9(2)	11.5(2)
²⁰⁸ Tl	583.1	350.7(16)	10.9(2)	7.8(2)
²²⁸ Ac	911.1	255.3(14)	1.1(1)	0.9(1)
²²⁸ Ac	968.9	151.5(11)	0.6(1)	0.6(1)
²¹⁴ Bi	1120.3	83.2(9)	1.1(1)	1.7(1)
²¹⁴ Bi	1238.1	38.3(12)	0.4(0)	0.7(1)
⁴⁰ K	1460.8	1753.6(3)	6.3(1)	5.9(1)

*) Characteristic X-rays

In addition, below 95 keV characteristic X-ray emitted from Pb and Bi was shown apparently after shield[2,3]. In order to reduce these characteristic X-rays, 3 mm copper plates were added inside of the shield. As a result, the Xray peaks have decreased. In the case of the characteristic X-rays of 74.8 keV emitted from 214 Pb, its count rate has decreased by a factor of 4 from $27.8(2) \times 10^{-3}$ cps to $6.9(2) \times 10^{-3}$ cps.

3. Comparison of spectra of standard sources

Standard sources, ²⁴¹Am, ¹³³Ba, ¹³⁷Cs, ⁶⁰Co, ¹⁵²Eu, were measured with and without the shield as well as with the copper plates inside of the shield for 36,000 s in each case. Figure 2 shows the spectra of ¹³⁷Cs.



Figure 2. Spectra of Cs-137 under various shield conditions.

It was noticeable that backscatter peak and Compton edge were shown because of the decreased background. After adding the copper plates, the count rates of characteristic X-rays have decreased by a factor 5 whereas backscatter peak has increased. It is thought that the γ -ray emitted from ¹³⁷Cs is backscattered on the surface of the copper plates and detected[4].

4. Conclusion

Analyses on the background and standard source spectra both with and without the shield were performed. Under adding the shield, most of the count rates of the background γ -rays have decreased by a factor of 100. In order to reduce the characteristic X-rays, copper plates with 3 mm thickness were added inside the lead bricks, which resulted in X-rays reduction effect. On the measurement of the standard sources, Compton edge and backscatter peak were observed due to the reduced background by shield. When the copper plates were added, the X-ray counts have decreased while the backscatter peak has increased.

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